# Viscous flow in a cylindrical tube containing a line of spherical particles 

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The viscous, creeping flow through a cylindrical tube of a liquid, which contains rigid, spherical particles, is investigated analytically. The spheres are located on the axis of the cylinder and are equally spaced. Solutions are derived for particles in motion and fixed, with and without fluid discharge. Numerical results are presented for the drag on each sphere and the mean pressure drop for a wide range of sizes and spacings of the spheres. The study is motivated by possible application to blood flow in capillaries, where red blood cells represent particles of the same order of magnitude as the diameter of the capillary itself. The results may also be of interest in other applications, such as sedimentation and fluidized beds. It is shown that there is little interaction between particles if the spacing is more than one tube diameter, and that the additional pressure drop over that for Poiseuille flow is less than $50 \%$ if the sphere diameter is less than 0.8 of the tube diameter.

## 1. Introduction

The present study was motivated by the possible application of the results to blood flow in capillary blood vessels, in which the red blood cells represent particles whose diameter is of the same order of magnitude as the diameter of the blood vessel itself. Red blood cells are flexible, biconcave disks, but to permit analytic treatment they are herein approximated by rigid spheres.

In the idealized problem considered (see figure 1), an infinite row of rigid spheres is assumed to be moving in a viscous, incompressible fluid bounded by a fixed circular cylinder. The spheres are equally spaced along the axis of the cylinder, and the Reynolds number is assumed to be so low that inertial terms may be neglected.

It can be shown that, for creeping flow, the fluid motion for the geometry shown in figure 1 is unique for a given velocity of the spheres, $U$, and a given discharge, $Q$ (Wang \& Skalak 1967). Solutions are derived below for arbitrary $U$ and $Q$, but three subcases are of interest: (i) $U=0, Q \neq 0$, which represents flow in a cylindrical tube containing spheres which are each fixed in position; (ii) $U \neq 0, Q=0$, which represents a line of spheres moving in a fluid otherwise at rest as in the case of particles settling in a cylinder of fluid; (iii) the case of

[^0]zero drag on each particle, which is achieved if the fluid and particles are being pumped steadily through the tube by a pressure gradient, as is the case in blood flow.

In cases (i) and (ii), the drag on each particle is different from zero, and there is also a mean pressure gradient along the cylinder. By superimposing the two cases of non-zero drag in proper proportion, a net drag equal to zero can be achieved. Numerical results are given below for all three cases.

Previous investigators have considered the influence of a cylindrical wall on the drag exerted on a single sphere (Haberman \& Sayre 1958), but no comparable results are available for an infinite line of spheres. A general discussion and some numerical results were given by Sonshine \& Brenner (1966). The present paper gives numerical results for pressure drop and drag for a range of the ratio, $\lambda$, of the sphere radius to the tube radius of $0<\lambda \leqslant 0 \cdot 9$, and a range of sphere spacing from the case of adjacent spheres touching up to a maximum spacing of forty tube radii. The present paper is complementary to Lighthill (1968), which considers elastic particles that nearly fit the tube, using lubrication theory.
The analytic techniques used in the present paper were suggested by the work of Ling (1963) and Atsumi (1960) on problems in the theory of elasticity involving similar geometry, e.g. torsion and tension of a circular bar containing an infinite row of spherical cavities.

## 2. Formulation

The flows of interest are assumed to be axisymmetric, so that a Stokes stream function exists. It is convenient to work in terms of dimensionless (unprimed) co-ordinates, which are defined in terms of the dimensional (primed) co-ordinates (figure l) by

$$
\begin{equation*}
R=R^{\prime}\left|a, \quad z=z^{\prime}\right| a, \quad r=r^{\prime} \mid a \tag{2.1}
\end{equation*}
$$



Figure 1. Geometry of the problem.
where $a$ is the radius of the cylinder, $\left(R^{\prime}, z^{\prime}\right)$ are cylindrical co-ordinates, and $\left(r^{\prime}, \theta\right)$ are spherical co-ordinates. We assume that the body force, if any, is a constant, $g$, per unit mass, and that it acts only in the axial direction, $+z$. If the
body force acts in a direction other than the axial direction, then the particles must be neutrally buoyant in order to remain concentric in the cylinder. The present analysis can be applied to this case by setting $g=0$.

The dependent variables of interest are the velocity, $\mathrm{V}^{\prime}$, pressure, $p^{\prime}$, stress tensor, $\tau_{i j}^{\prime}$, Stokes stream function, $\psi^{\prime}$, drag $D^{\prime}$, and discharge, $Q^{\prime}$. These are expressed in dimensionless form, using the fluid density and kinematic viscosity, $\rho$ and $\nu$, and the tube radius, $a$, as follows:

$$
\left.\begin{array}{l}
\mathbf{V}=a \mathbf{V}^{\prime} / v, \quad p=\frac{p^{\prime} a^{2}}{\rho \nu^{2}}-\frac{a^{2} g z^{\prime}}{\nu^{2}}, \quad \tau_{i j}=\tau_{i j}^{\prime} a^{2} / \rho \nu^{2},  \tag{2.2}\\
D=D^{\prime} \mid \rho \nu^{2}, \quad \psi=\psi^{\prime} / a v, \quad Q=Q^{\prime} / a v .
\end{array}\right\}
$$

In (2.2) the effect of the body force is incorporated in the dimensionless pressure. Then the equations of motion and continuity are

$$
\begin{gather*}
-\nabla p+\nabla^{2} \mathbf{V}=0  \tag{2.3}\\
\nabla . \mathbf{V}=0 \tag{2.4}
\end{gather*}
$$

where $\nabla$ is the gradient operator in dimensionless co-ordinates. Taking the curl of (2.3) gives

$$
\begin{equation*}
\nabla^{2} \zeta=0 \tag{2.5}
\end{equation*}
$$

where $\zeta$ is the dimensionless vorticity, $\nabla \times V$. Equation (2.5) yields a fourthorder equation, $L^{4} \psi=0$, for axisymmetric flow. In cylindrical co-ordinates (Happel \& Brenner 1965),

$$
\begin{equation*}
\left[\frac{\partial^{2}}{\partial z^{2}}+\frac{\partial^{2}}{\partial R^{2}}-\frac{1}{R} \frac{\partial}{\partial R}\right]^{2} \psi(z, R)=0 \tag{2.6}
\end{equation*}
$$

and the velocity components are

$$
\begin{equation*}
v_{z}=-\frac{1}{R} \frac{\partial \psi}{\partial R}, \quad v_{R}=\frac{1}{R} \frac{\partial \psi}{\partial z} . \tag{2.7}
\end{equation*}
$$

In spherical co-ordinates, $L^{4} \psi=0$ is (Happel \& Brenner 1965)

$$
\begin{equation*}
\left[\frac{\partial^{2}}{\partial r^{2}}+\frac{1-\cos ^{2} \theta}{r^{2}} \frac{\partial^{2}}{\partial(\cos \theta)^{2}}\right]^{2} \psi(r, \theta)=0 \tag{2.8}
\end{equation*}
$$

and the velocity components are

$$
\begin{equation*}
v_{r}=-\frac{1}{r^{2} \sin \theta} \frac{\partial \psi}{\partial \theta}, \quad v_{\theta}=\frac{1}{r \sin \theta} \frac{\partial \psi}{\partial r} . \tag{2.9}
\end{equation*}
$$

The general problem considered is to find the stream function, $\psi$, for arbitrary sphere velocity, $U$, and discharge, $Q$ (figure 1). The discharge, $Q$, is computed on any cross-section of the tube and includes both the solid and liquid components. In terms of the stream function, $Q$ is given by

$$
\begin{equation*}
Q=-2 \pi[\psi]_{R=1}, \tag{2.10}
\end{equation*}
$$

where it assumed that $\psi$ is adjusted so that $[\psi]_{R=0}=0$.

The boundary conditions to be satisfied by $\psi$ are as follows: at the cylinder walls, $R=1$,

$$
\begin{equation*}
\psi=-\frac{Q}{2 \psi} \quad v_{z}=0 \tag{2.11}
\end{equation*}
$$

on the spherical surfaces, $r_{n}=\lambda, n=0, \pm 1, \pm 2, \ldots$,

$$
\begin{gather*}
v_{z}=U, \quad v_{R}=0,  \tag{2.12}\\
r_{n}=\left[R^{2}+(z-n \beta)^{2}\right]^{\frac{1}{2}}, \tag{2.13}
\end{gather*}
$$

where
and $\lambda$ and $\beta$ are the dimensionless radius and spacing of the spheres (figure 1 ). The above boundary conditions apply in a co-ordinate system fixed with respect to the tube wall. It is convenient to use instead a co-ordinate system fixed with respect to the spheres in which the boundary conditions are as follows:
at the cylinder walls, $R=1$,

$$
\begin{equation*}
\psi=-\frac{Q}{2 \pi}+\frac{U}{2}, \quad v_{z}=-U \tag{2.14}
\end{equation*}
$$

on the spherical surfaces, $r_{n}=\lambda$,

$$
\begin{equation*}
\psi=0, \quad \frac{\partial \psi}{\partial r_{n}}=0 \tag{2.15}
\end{equation*}
$$

In (2.14) and below, $Q$ and $U$ are the discharge and sphere velocity in co-ordinates fixed with respect to the tube wall. The stream function, $\psi$, is written for coordinates moving with the spheres. Now let

$$
\begin{equation*}
\psi=\psi_{0}+\psi_{a} \tag{2.16}
\end{equation*}
$$

where $\psi_{0}$ is the Poiseuille flow that would take place if there were no spheres present;

$$
\begin{gather*}
\psi_{0}=\left(\frac{U}{2}-\frac{V}{2}\right) R^{2}+\frac{V}{4} R^{4}  \tag{2.17}\\
V=\frac{2 Q}{\pi} \tag{2.18}
\end{gather*}
$$

The boundary conditions on the additional stream function, $\psi_{a}$, are then:
at the cylinder walls, $R=1$,

$$
\begin{equation*}
v_{z}=-\frac{1}{R} \frac{\partial \psi_{a}}{\partial R}=0, \quad \psi_{a}=0 \tag{2.19}
\end{equation*}
$$

on the spherical surfaces, $r_{n}=\lambda$,

$$
\begin{equation*}
\psi_{a}=-\psi_{0}, \quad \frac{\partial \psi_{a}}{\partial r_{n}}=-\frac{\partial \psi_{0}}{\partial r_{n}} \tag{2.20}
\end{equation*}
$$

The problem is now to find a $\psi_{a}$ which satisfies $L^{4} \psi_{a}=0$ and the boundary conditions (2.19) and (2.20).

## 3. Solution for $\psi$

Solutions to $L^{4} \psi=0$ which are useful in the present case are given by Haberman \& Sayre (1958). Solutions in cylindrical co-ordinates, which are even and periodic in $z$, with period $\beta$, and which give finite velocities everywhere, are

$$
\begin{equation*}
\psi(z, R)=A_{0} R^{4}+B_{0} R^{2}+\sum_{m=1}^{\infty}\left[A_{m} R I_{1}(m K R)+B_{m} R^{2} I_{0}(m K R)\right] \cos m K z \tag{3.1}
\end{equation*}
$$

where $K=2 \pi / \beta$ and $I_{1}$ and $I_{0}$ are modified Bessel functions of the first kind. $A_{m}, B_{m}$ are arbitrary constants.

Solutions in spherical co-ordinates, which are even in $z$ and give rise to finite velocities everywhere, except possibly at infinity and at the origin, are

$$
\begin{equation*}
\psi(r, \theta)=\sum_{n=2}^{\infty} C_{n}^{-\frac{1}{2}}(\cos \theta)\left[A_{n} r^{n}+B_{n} r^{-n+1}+C_{n} r^{n+2}+D_{n} r^{-n+3}\right], \tag{3.2}
\end{equation*}
$$

where $n=2,4,6, \ldots$, and $C_{n}^{-\frac{1}{2}}(\cos \theta)$ are Gegenbauer functions of order $n$ and degree $-\frac{1}{2} . A_{n}, B_{n}, C_{n}, D_{n}$ are arbitrary constants. The present problem requires solutions which are periodic as well as even in $z$. Periodic functions in spherical co-ordinates can be constructed by replacing $z$ in (3.2) by $(z-n \beta)$ and summing over $n$ from $-\infty$ to $+\infty$. Physically, this corresponds to placing appropriate singularities at the centre of each sphere in figure 1. The following two even, periodic solutions are constructed in this way:

$$
\begin{align*}
\psi^{(1)}= & \sum_{n=-\infty}^{\infty} \frac{R^{2}}{2\left[R^{2}+(z-n \beta)^{2}\right]^{\frac{3}{2}}}+\left\{A_{0} R^{4}+B_{0} R^{2}+\sum_{m=1}^{\infty}\left[A_{m} R I_{1}(m K R)\right.\right. \\
& \left.\left.+B_{m} R^{2} I_{0}(m K R)\right] \cos m K z\right\},  \tag{3.3}\\
\Psi^{(2)}= & \left\{\frac{1}{2} \frac{R^{2}}{\left[R^{2}+z^{2}\right]^{\frac{1}{2}}}+\frac{1}{2} \sum_{n=1}^{\infty}\left[\frac{R^{2}}{\left[R^{2}+(z-n \beta)^{2}\right]^{\frac{1}{2}}}+\frac{R^{2}}{\left[R^{2}+(z+n \beta)^{2}\right]^{\frac{1}{2}}}-\frac{2 R^{2}}{n \beta}\right]\right\} \\
& +\left\{C_{0} R^{4}+D_{0} R^{2}+\sum_{m=1}^{\infty}\left[C_{m} R I_{1}(m K R)+D_{m} R^{2} I_{0}(m K R)\right] \cos m K z\right\}, \tag{3.4}
\end{align*}
$$

where $A_{m}, B_{m}, C_{m}$, and $D_{m}$ are arbitrary constants. The portions of the solutions containing these constants will be used to satisfy the boundary conditions on the cylindrical wall.

The first term of $\psi^{(1)}$ comes from $C_{2}^{-\frac{1}{2}}(\cos \theta) r^{-2+1}=R^{2} / 2 r^{3}$, and the first term of $\psi^{(2)}$ comes from $C_{2}^{-\frac{1}{2}}(\cos \theta) r^{-2+3}=R^{2} / 2 r$. These two terms may be regarded as the terms which occur in the Stokes solution for a single sphere in an infinite fluid.

The functions $\psi^{(1)}$ and $\psi^{(2)}$ are now adjusted so that each satisfies the boundary conditions at the cylindrical wall,

$$
\begin{equation*}
\psi \psi=v_{z}=0 \quad \text { on } \quad R=1 \tag{3.5}
\end{equation*}
$$

For $\psi^{(1)}$, the boundary conditions (3.5) require:
for

$$
\begin{align*}
\psi & =0: A_{0}+B_{0}+\sum_{m=1}^{\infty}\left[A_{m} I_{1}(m K)+B_{m} I_{0}(m K)\right] \cos m K z, \\
& =-\sum_{n=-\infty}^{\infty} \frac{1}{2\left[1+(z-n \beta)^{2}\right]^{\frac{3}{2}}} \tag{3.6}
\end{align*}
$$

for $\quad v_{z}=0: 4 A_{0}+2 B_{0}+\sum_{m=1}^{\infty}\left[A_{m} m K I_{0}(m K)\right.$

$$
\left.+B_{m}\left(2 I_{0}(m K)+m K I_{1}(m K)\right)\right] \cos m K z
$$

$$
\begin{equation*}
=-\left\{\sum_{n=-\infty}^{\infty} \frac{1}{\left[1+(z-n \beta)^{2}\right]^{\frac{2}{2}}}-\frac{3}{2} \frac{1}{\left[1+(z-n \beta)^{2}\right]^{\frac{3}{2}}}\right\} . \tag{3.7}
\end{equation*}
$$

Both sides of (3.6) and (3.7) are multiplied by $\cos m K z$ and then integrated from $-\frac{1}{2} \beta$ to $\frac{1}{2} \beta$, to derive a pair of algebraic equations for each pair of the $A_{m}, B_{m}$, $m=0,1,2,3, \ldots$ A similar set of equations for each of the pairs $C_{m}, D_{m}$ is derived by applying (3.5) to $\psi^{(2)}$. The following results are obtained by solving the several sets simultaneously:

$$
\left.\begin{array}{rl}
A_{m}= & \frac{2}{\beta} \frac{(m K)^{2}\left[I_{1}(m K) K_{1}(m K)+I_{0}(m K) K_{2}(m K)\right]}{m K\left[I_{0}^{2}(m K)-I_{1}^{2}(m K)\right]-2 I_{0}(m K) I_{1}(m K)}, \\
B_{m}= & \left.\frac{2}{\beta} \frac{(m K)^{2}\left[-I_{0}(m K) K_{1}(m K)-I_{1}(m K) K_{2}(m K)\right]+2 m K I_{1}(m K) K_{1}(m K)}{m K\left[I_{0}^{2}(m K)-I_{1}^{2}(m K)\right]-2 I_{0}(m K) I_{1}(m K)},\right\} \\
A_{0}= & \frac{1}{\beta}, \quad B_{0}=\frac{-2}{\beta}, \quad C_{m}=-\frac{B_{m}}{m K} ; \\
D_{m}= & \frac{2}{\beta}\left\{(m K)\left[-I_{0}(m K) K_{2}(m K)-I_{1}(m K) K_{1}(m K)\right]+\left[2 I_{0}(m K) K_{1}(m K)\right.\right. \\
& \left.\left.\left.\left.+2 K_{0}(m K) I_{1}(m K)\right]\right\} /\left\{(m K)\left[I_{0}^{2}(m K)-I_{1}^{2}(m K)\right]-2 I_{0}[m K) I_{1}(m K)\right]\right\},\right\}  \tag{3.9}\\
C_{0}= & \frac{1}{2 \beta}, \quad D_{0}=\frac{1}{2 \beta}\{-2 \log (2 \beta)+2 E-1\},
\end{array}\right\}
$$

where Euler's constant

$$
\begin{equation*}
E=\lim _{N \rightarrow \infty}\left[\sum_{n=1}^{N} \frac{1}{n}-\log N\right]=0.577215665 \tag{3.10}
\end{equation*}
$$

The modified Bessel functions of the second kind, $K_{n}$, which occur in (3.8) and (3.9) arise from the integrals with respect to $z$ (see Watson 1944, p. 185).

Next, we note that even derivatives of $\psi^{(1)}$ and $\psi^{(2)}$ with respect to $z$ generate additional functions, each of which satisfies $L^{4} \psi=0$ and the boundary conditions (3.5) at $R=1$. They are also periodic and even in $z$. Hence, we define

$$
\begin{equation*}
\psi_{2 s}^{(i)}=\frac{1}{(2 s-2)!} \frac{\partial^{2 s-2} \psi^{(i)}}{\partial z^{2 s-2}} \quad(s \geqslant 1, i=1,2) . \tag{3.11}
\end{equation*}
$$

The differentiations called for in (3.11) generate a complete set of functions, which contain singularities of all the even orders $n$ in (3.2). The stream function $\psi_{a}$ is assumed to be of the form

$$
\begin{equation*}
\psi_{a}=\sum_{s=1}^{\infty}\left[E_{2 s} \psi_{2 s}^{(1)}+F_{2 s} \psi_{2 s}^{(2)}\right] \tag{3.12}
\end{equation*}
$$

The constants $E_{2 s}$ and $F_{2 s}$ in (3.12) are to be determined so that the boundary conditions (2.20) on the spherical surfaces are satisfied. For this purpose, the functions defined by (3.11) are expressed in spherical co-ordinates. This is a critical and important step of the entire procedure. As an example of the procedure, consider the first series in (3.4), which may be written in the form

$$
\begin{align*}
& \sum_{n=-\infty}^{\infty} \frac{R^{2}}{2\left[R^{2}+(z-n \beta)^{2}\right]^{\frac{3}{2}}}=\frac{R^{2}}{2 r^{3}}+\sum_{n=1}^{\infty} \frac{R^{2}}{2(n \beta)^{3}\left[1+2 h_{n} \mu+h_{n}^{2}\right]^{\frac{3}{2}}} \\
&+\sum_{n=1}^{\infty} \frac{R^{2}}{2(n \beta)^{3}\left[1-2 h_{n} \mu+h_{n}^{2}\right]^{\frac{3}{2}}}, \tag{3.13}
\end{align*}
$$

where $\mu=\cos \theta$ and $h_{n}=r / n \beta$. The sums on the right-hand side of (3.13) are recognized in terms of Legendre functions using the relations (Hobson 1955, pp. 105, 107):

$$
\begin{gather*}
\frac{1.3 .5 \ldots(2 m-1) h^{m}}{\left(1-2 h \mu+h^{2}\right)^{m+\frac{1}{2}}}=\sum_{n=m}^{\infty} h^{n} \frac{d^{m} P_{n}(\mu)}{d \mu^{m}}  \tag{3.14}\\
(-1)^{m}\left(1-\mu^{2}\right)^{\frac{1}{2} m} \frac{d^{m} P_{n}(\mu)}{d \mu^{m}}=P_{n}^{m}(\mu) \tag{3.15}
\end{gather*}
$$

where $P_{n}$ are Legendre polynomials of $n$th degree, and $P_{n}^{m}$ are associated Legendre functions of the $m$ th order and $n$th degree. Using (3.14) and (3.15), equation (3.13) becomes

$$
\begin{align*}
& \sum_{n=-\infty}^{\infty} \frac{R^{2}}{2\left[R^{2}+(z-n \beta)^{2}\right]^{\frac{3}{2}}}=\frac{R^{2}}{2 r^{3}}-\frac{1}{2} R \sum_{n=1}^{\infty} \sum_{q=0}^{\infty} \frac{(-1)^{q} q^{q+1}}{(n \beta)^{q+3}} P_{1+q}^{1}(\cos \theta) \\
&-\frac{1}{2} R \sum_{n=1}^{\infty} \sum_{q=0}^{\infty} \frac{r^{q+1}}{(n \beta)^{q+3}} P_{1+q}^{1}(\cos \theta) . \tag{3.16}
\end{align*}
$$

The derivatives with respect to $z$ of (3.16) are also required, and they are found using known recurrence relations (Hobson 1955; pp. 105, 138). The end result is

$$
\begin{align*}
& \frac{\partial^{2 s-2}}{\partial z^{2 s-2}} \sum_{n=-\infty}^{\infty} \frac{R^{2}}{2\left[R^{2}+(z-n \beta)^{2}\right]^{\frac{3}{2}}} \\
&=-\frac{R}{2} \frac{(2 s-2)!}{r^{2 s}} P_{2 s-1}^{1}(\cos \theta) \\
& \quad-R \sum_{n=1}^{\infty} \sum_{l=0}^{\infty} \frac{1}{(n \beta)^{2 l+2 s+1}} \frac{(2 l+2 s)!}{(2 l+2)!} r^{2 l+1} P_{2 l+1}^{1}(\cos \theta) . \tag{3.17}
\end{align*}
$$

The terms containing $I_{0}$ and $I_{1}$ in (3.3) and (3.4) must also be converted to spherical co-ordinates. The key relation used for this purpose is (MacRobert 1948, p. 109)

$$
\begin{equation*}
\sum_{n=0}^{\infty} \frac{R^{n}}{n!} P_{n}(\cos \theta)=e^{x} J_{0}(y) \tag{3.18}
\end{equation*}
$$

where $x=R \cos \theta$ and $y=R \sin \theta$. Equation (3.18) and the two relations

$$
\begin{gather*}
I_{\nu}(z)=e^{-\nu \pi i / 2} J_{\nu}(i z),  \tag{3.19}\\
\sin \theta P_{2 n}=-\frac{P_{2 n+1}^{1}}{4 n+1}+\frac{P_{2 n-1}^{1}}{4 n+1} \tag{3.20}
\end{gather*}
$$

are used to arrive at

$$
\begin{align*}
& \frac{\partial^{2 s-2}}{\partial z^{2 s-2}}\left[R^{2} I_{0}(m K R) \cos m K z\right] \\
& \quad=R \sum_{l=0}^{\infty}\left\{\frac{(-1)^{l+s}(m K)^{2 l+2 s-2}}{(2 l+2)!} r^{2 l+1}\left[\frac{(2 l+2)(2 l+1)}{4 l+1}+\frac{(m K)^{2} r^{2}}{4 l+5}\right] P_{2 l+1}^{1}\right\} . \tag{3.21}
\end{align*}
$$

Since $I_{1}(x)=(d / d x) I_{0}(x)$, equation (3.21) can also be used to derive

$$
\begin{equation*}
\frac{\partial^{2 s-2}}{\partial z^{2 s-2}}\left[R I_{1}(m K R) \cos m K z\right]=R \sum_{l=0}^{\infty} \frac{(-1)^{l+s}(m K)^{2 l+2 s-1}}{(2 l+2)!} r^{2 l+1} P_{2 l+1}^{1} . \tag{3.22}
\end{equation*}
$$

The sum of (3.17), (3.21) and (3.22) expresses $\psi_{2 \mathrm{~s}}^{(1)}$ in spherical co-ordinates. Similar procedures are used to convert $\psi_{2 s}^{(2)}$ to spherical co-ordinates. Finally, $\psi_{0}$, equation (2.17), is expressed in spherical co-ordinates as

$$
\begin{equation*}
\psi_{0}=R\left\{P_{1}^{1}\left[\left(-\frac{U}{2}+\frac{V}{2}\right) r-\frac{V r^{3}}{5}\right]+P_{3}^{1} \frac{V r^{3}}{30}\right\} . \tag{3.23}
\end{equation*}
$$

In spherical co-ordinates, the complete stream function, $\psi$, is

$$
\begin{align*}
\psi=\psi_{a}+\psi_{0}=\sin \theta[ & \sum_{s=1}^{\infty} F_{2 s}\left\{\frac{1}{2(4 s-3)} \frac{1}{r^{2 s-3}}\left(P_{2 s-3}^{1}-P_{2 s-1}^{1}\right)\right. \\
& +\delta_{1 s}\left[-\frac{r^{2}}{2 \beta}(-2 \log (2 \beta)+2 E-1) P_{1}^{1}+\frac{r^{4}}{2 \beta}\left(-\frac{4}{5} P_{1}^{1}+\frac{4}{30} P_{3}^{1}\right)\right] \\
& \left.+\sum_{l=0}^{\infty}\left[r^{2 l+2}\left(-\tau_{2 s}^{2 l}+\left(\beta_{2}\right)_{2 s}^{2 l}+\left(\gamma_{2}\right)_{2 s}^{2 l}\right)+r^{2 l+4}\left(\xi_{2 s}^{2 l}+\left(\sigma_{2}\right)_{2 s}^{2 l}\right)\right] P_{2 l+1}^{1}\right\} \\
& +\sum_{s=1}^{\infty} E_{2 s}\left\{\frac{-1}{2 r^{2 s-1}} P_{2 s-1}^{1}+\delta_{1 s}\left[\frac{2}{\beta} r^{2} P_{1}^{1}+\frac{1}{\beta} r^{4}\left(-\frac{4}{5} P_{1}^{1}+\frac{4}{30} P_{3}^{1}\right)\right]\right. \\
& \left.+\sum_{l=0}^{\infty}\left[r^{2 l+2}\left(-\alpha_{2 s}^{2 l}+\left(\beta_{1}\right) \frac{2 s}{2 l}+\left(\gamma_{1}\right)_{2 s}^{2 l}\right)+r^{2 l+4}\left(\sigma_{1}\right)_{2 s}^{2 l}\right] P_{22+1}^{1}\right\} \\
& \left.+\left\{P_{1}^{1}\left[\left(\frac{V}{2}-\frac{U}{2}\right) r^{2}-\frac{V r^{4}}{5}\right]+P_{3}^{1} \frac{V r^{4}}{30}\right\}\right] \tag{3.24}
\end{align*}
$$

where

$$
\left.\begin{array}{rl}
\tau_{2 s}^{2 l} & =\sum_{n=1}^{\infty} \frac{1}{(n \beta)^{2 l+2 s-1}} \frac{(2 l+2 s-2)!}{(2 s-2)!(2 l)!} \frac{1}{4 l+1}, \\
\xi_{2 s}^{2 l} & =\sum_{n=1}^{\infty} \frac{1}{(n \beta)^{2 l+2 s+1}} \frac{(2 l+2 s)!}{(2 s-2)!(2 l+2)!} \frac{1}{4 l+5}, \\
\left(\beta_{2}\right)_{2 s}^{2 l} & =\sum_{m=1}^{\infty} C_{m} \frac{(-1)^{l+s}(m K)^{2 l+2 s-1}}{(2 l+2)!(2 s-2)!}, \\
\left(\gamma_{2}\right)_{2 s}^{l l} & =\sum_{m=1}^{\infty} D_{m} \frac{(-1)^{l+s}(m K)^{2 l+2 s-2}}{(2 l)!(2 s-2)!(4 l+1)},  \tag{3.25}\\
\left(\sigma_{2}\right)_{2 s}^{2 l} & =\sum_{m=1}^{\infty} D_{m} \frac{(-1)^{l+s}(m K)^{2 l+2 s}}{(2 l+2)!(2 s-2)!(4 l+5)}, \\
\alpha_{2 s}^{2 l} & =\sum_{n=1}^{\infty} \frac{1}{(n \beta)^{2 l+2 s+1}} \frac{(2 l+2 s)!}{(2 l+2)!(2 s-2)!}, \\
\left(\beta_{1}\right)_{2 s}^{2 l} & =\sum_{m=1}^{\infty} A_{m} \frac{(-1)^{l+s}(m K)^{2 l+2 s-1}}{(2 l+2)!(2 s-2)!}, \\
\left(\gamma_{1}\right)_{2 s}^{l l} & =\sum_{m=1}^{\infty} B_{m} \frac{(-1)^{l+s}(m K)^{2 l+2 s-2}}{(2 l)!(2 s-2)!(4 l+1)}, \\
\left(\sigma_{1}\right)_{2 s}^{2 l} & =\sum_{m=1}^{\infty} B_{m} \frac{(-1)^{l+s}(m K)^{2 l+2 s}}{(2 l+2)!(2 s-2)!(4 l+5)} .
\end{array}\right\}
$$

Each of the formulas in (3.25) holds for $s=1,2,3, \ldots$ and $l=0,1,2, \ldots$ with the one exception that for $s=1$ and $l=0$

$$
\begin{equation*}
\tau_{2 s}^{2 l}=0 \tag{3.26}
\end{equation*}
$$

It remains to satisfy the boundary conditions (2.15) on the surface of each sphere. Since $\psi$ is periodic in $z$, it suffices to satisfy (2.15) on the sphere centred at the origin. This leads to two equations, each of which is a linear sum of the functions $P_{2 l+1}^{1}(l=0,1,2, \ldots)$. Since the $P_{2 l+1}^{1}$ form a linearly independent set, the zero boundary conditions (2.15) require the coefficient of each $P_{2 l+1}^{1}$ to be zero. This yields two infinite sets of linear, algebraic equations on the coefficients $E_{2 s}$ and $F_{2 s}$, which may be written in the form

$$
\begin{align*}
&-\frac{F_{2 l+2}}{4 l+1}+F_{2}\left\{\left[-\frac{3 \lambda}{2 \beta}(-2 \log (2 \beta)+2 E-1)-\frac{2 \lambda^{3}}{\beta}\right] \delta_{0 l}+\frac{28}{60 \beta} \lambda^{5} \delta_{1 l}\right\} \\
&+ E_{2}\left\{\left(\frac{6 \lambda}{\beta}-\frac{4 \lambda^{3}}{\beta}\right) \delta_{0 l}+\frac{28}{30 \beta} \lambda^{5} \delta_{1 l}\right\} \\
&++\sum_{s=1}^{\infty}\left\{F_{2 s}\left[(4 l+3)\left(-\tau_{2 s}^{2 l}+\left(\beta_{2}\right)_{2 s}^{2 l}+\left(\gamma_{2}\right)_{2 s}^{2 l}\right) \lambda^{4 l+1}+(4 l+5)\left(\xi_{2 s}^{2 l}+\left(\sigma_{2}\right)_{2 s}^{2 l}\right) \lambda^{4 l+3}\right]\right. \\
&+\left.E_{2 s}\left[(4 l+3)\left(-\alpha_{2 s}^{2 l}+\left(\beta_{1}\right)_{2 s}^{2 l}+\left(\gamma_{1}\right)_{2 s}^{2 l}\right) \lambda^{4 l+1}+(4 l+5)\left(\sigma_{1}\right)_{2 s}^{2 l} \lambda^{4 l+3}\right]\right\} \\
&=U \frac{3}{2} \lambda \delta_{0 l}+V\left\{\left(-\frac{3}{2} \lambda+\lambda^{3}\right) \delta_{0 l}-\frac{7}{30} \lambda^{5} \delta_{1 l}\right\},  \tag{3.27}\\
&-\frac{F_{2 l+4}}{4 l+5}+E_{2 l+2}+F_{2}\left\{\left(-\frac{\lambda^{3}}{2 \beta}(-2 \log (2 \beta)+2 E-1)-\frac{12}{10} \frac{\lambda^{5}}{\beta}\right] \delta_{0 l}+\frac{1}{3 \beta} \lambda^{7} \delta_{1 l}\right\} \\
&+ E_{2}\left\{\left(\frac{2 \lambda^{3}}{\beta}-\frac{12}{5 \beta} \lambda^{5}\right) \delta_{0 l}+\frac{2}{3 \beta} \lambda^{7} \delta_{1 l}\right\} \\
&+ \sum_{s=1}^{\infty}\left\{F_{2 s}\left[(4 l+1)\left(-\tau_{2 s}^{2 l}+\left(\beta_{2}\right)_{2 s}^{2 l}+\left(\gamma_{2}\right)_{2 s}^{2 l}\right) \lambda^{4 l+3}+(4 l+3)\left(\xi_{2 s}^{2 l}+\left(\sigma_{2}\right)_{2 s}^{2 l}\right) \lambda^{4 l+5}\right]\right. \\
&+\left.E_{2 s}\left[(4 l+1)\left(-\alpha_{2 s}^{2 l}+\left(\beta_{1}\right)_{2 s}^{2 l}+\left(\gamma_{1}\right)_{2 s}^{2 l}\right) \lambda^{4 l+3}+(4 l+3)\left(\sigma_{1}\right)_{2 S}^{2 l} \lambda^{4 l+5}\right]\right\} \\
&=\frac{U}{2} \lambda^{3} \delta_{0 l}+V\left\{\left(-\frac{1}{2} \lambda^{3}+\frac{3 \lambda^{5}}{5}\right) \delta_{0 l}-\frac{5 \lambda^{7}}{30} \delta_{1 l}\right\} . \tag{3.28}
\end{align*}
$$

The two sets of equations (3.27) and (3.28) were solved simultaneously by successive approximations. Both sets were truncated, so that the combined sets contained the same number of equations as unknowns starting from a $2 \times 2$ matrix up to a maximum of $16 \times 16$.

## 4. Drag and pressure drop

Once the coefficients $E_{2 s}$ and $F_{2 s}$ are computed, the stream function is completely determined, and the physical quantities of interest, such as drag and pressure drop, can be computed. Since each sphere is moving at a constant velocity, the equation of motion requires that the sum of all the forces on each sphere be zero. For the typical sphere centred at the origin, the sum of the forces, $F_{t}$, in dimensionless form, is

$$
\begin{equation*}
F_{t}=F_{e}+W_{s}+2 \pi \lambda^{2} \int_{0}^{\pi}\left(\tau_{r r} \cos \theta-\tau_{r \theta} \sin \theta\right) \sin \theta d \theta=0 \tag{4.1}
\end{equation*}
$$

where $F_{e}=F_{e}^{\prime} / \rho \nu^{2}$ and $F_{e}^{\prime}$ is the applied external force on each particle. $W_{s}=W_{s}^{\prime} / \rho \nu^{2}$ and $W_{s}^{\prime}$ is the weight of the particle. The integral term in (4.1)
includes both the drag, $D$, due to the viscosity of the fluid, and the buoyant effect of the fluid. The stress components in (4.1) are

$$
\begin{align*}
& \tau_{r r}=-p+2 \frac{\partial v_{r}}{\partial r}+g z \frac{a^{3}}{\nu^{2}}  \tag{4.2}\\
& \tau_{r \theta}=\frac{\mathbf{1}}{r} \frac{\partial v_{r}}{\partial \theta}-\frac{v_{\theta}}{r}+\frac{\partial v_{\theta}}{\partial r} \tag{4.3}
\end{align*}
$$

$\tau_{r \theta}$ is evaluated from (4.3), (3.24) and (2.9), and recurrence relations:

$$
\begin{align*}
\tau_{r \theta}=\sum_{s=1}^{\infty} F_{2 s}\{ & -\frac{2(2 s)(s-1)}{(4 s-3) r^{2 s}} P_{2 s-1}^{1}+\frac{(2 s-1)(2 s-3)}{(4 s-3) r^{2 s}} P_{2 s-3}^{1} \\
& +\sum_{l=0}^{\infty}\left[2(2 l)(2 l+2)\left(-\tau_{2 s}^{2 l}+\left(\beta_{2}\right) \frac{2 l}{2 l}+\left(\gamma_{2}\right)_{2 s}^{2 l}\right) r^{2 l-1}\right. \\
& \left.\left.+2(2 l+1)(2 l+3)\left(\xi_{2 s}^{2 l}+\left(\sigma_{2}\right)_{2 s}^{2 l}\right) r^{2 l+1}\right] P_{2 l+1}^{1}\right\} \\
& +F_{2} \frac{4 r}{\beta}\left(-\frac{3}{5} P_{1}^{1}+\frac{4}{15} P_{3}^{1}\right)+E_{2} \frac{8 r}{\beta}\left(-\frac{3}{5} P_{1}^{1}+\frac{4}{15} P_{3}^{1}\right) \\
& +\sum_{s=1}^{\infty} E_{2 s}\left\{-\frac{(2 s+1)(2 s-1)}{r^{2 s+2}} P_{2 s-1}^{1}\right. \\
& +\sum_{l=0}^{\infty}\left[2(2 l)(2 l+2)\left(-\alpha_{2 s}^{2 l}+\left(\beta_{1}\right)_{2 s}^{2 l}+\left(\gamma_{1}\right) \frac{2 l}{2 l}\right) r^{2 l-1}\right. \\
& \left.\left.+2(2 l+1)(2 l+3)\left(\sigma_{1}\right)_{2 s}^{2 l} r^{2 l+1}\right] P_{2 l+1}^{1}\right\}+2 V r\left(-\frac{3}{5} P_{1}^{1}+\frac{4}{15} P_{3}^{1}\right) \tag{4.4}
\end{align*}
$$

The pressure is related to $\psi$ by (Happel \& Brenner 1965, p. 121)

$$
\begin{equation*}
\frac{\partial p}{\partial r}=-\frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \theta}\left(E^{2} \psi\right) \tag{4.5}
\end{equation*}
$$

where

$$
\begin{equation*}
E^{2} \psi=\frac{\partial^{2} \psi}{\partial r^{2}}+\frac{\sin \theta}{r^{2}} \frac{\partial}{\partial \theta}\left(\frac{1}{\sin \theta} \frac{\partial \psi}{\partial \theta}\right) \tag{4.6}
\end{equation*}
$$

After substituting (3.24) in (4.7), the pressure is obtained by integrating both (4.5) and (4.6). Comparing the two results, we find:

$$
\begin{align*}
p=\sum_{s=1}^{\infty} F_{2 s}\{ & \left.-\frac{2 s-1}{r^{2 s}} P_{2 s-1}+\sum_{l=0}^{\infty}(4 l+4)(4 l+5)\left(\xi_{2 s}^{2 l}+\left(\sigma_{2}\right)_{2 s}^{2 l}\right) r^{2 l+1} P_{2 l+1}\right\} \\
& +\sum_{s=1}^{\infty} E_{2 s}\left\{\sum_{l=0}^{\infty}(4 l+4)(4 l+5)\left(\sigma_{1}\right)_{2 s}^{2 l} r^{2 l+1} P_{2 l+1}\right\} \\
& -8 F_{2} \frac{r}{\beta} P_{1}-16 E_{2} \frac{r}{\beta} P_{1}-4 V r P_{1}+p_{0} \tag{4.8}
\end{align*}
$$

where $p_{0}$ is an arbitrary constant pressure. The normal stress (4.3) is then

$$
\begin{align*}
& \tau_{r r}=\sum_{s=1}^{\infty} F_{2 s}\left\{\frac{2 s-1}{4 s-3} \frac{(s-1)(4 s+6)+3}{r^{2 s}} P_{2 s-1}-\frac{(2 s-1)(2 s-2)(2 s-3)}{(4 s-3) r^{2 s}} P_{2 s-3}\right. \\
&+\sum_{l=0}^{\infty}\left[2(2 l+2)(2 l+1)(2 l)\left(-\tau_{2 s}^{2 l}+\left(\beta_{2}\right)_{2 s}^{2 l}+\left(\gamma_{2}\right)_{2 s}^{2 l}\right) r^{2 l-1}\right. \\
&\left.\left.+2(2 l+2)\left(4 l^{2}+2 l-3\right)\left(\xi_{2 s}^{2 l}+\left(\sigma_{2}\right)_{2 s}^{2 l}\right) r^{2 l+1}\right] P_{2 l+1}\right\} \\
&+\sum_{s=1}^{\infty} E_{2 s}\left(\frac{(2 s+1)(2 s)(2 s-1)}{r^{2 s+2}} P_{2 s-1}\right. \\
&+\sum_{l=0}^{\infty}\left[2(2 l+2)(2 l+1)(2 l)\left(-\alpha_{2 s}^{2 l}+\left(\beta_{1}\right)_{2 s}^{2 l}+\left(\gamma_{1}\right)_{2 s}^{2 l}\right) r^{2 l-1}\right. \\
&\left.\left.+2(2 l+2)\left(4 l^{2}+2 l-3\right)\left(\sigma_{1}\right)_{2 s}^{2 l} r^{2 l+1}\right] P_{2 l+1}\right\}+\frac{8 F_{2} r}{\beta}\left(\frac{3}{5} P_{1}+\frac{2}{5} P_{3}\right) \\
&+\frac{16 E_{2} r}{\beta}\left(\frac{3}{5} P_{1}+\frac{2}{5} P_{3}\right)+4 V r\left(\frac{3}{5} P_{1}+\frac{2}{5} P_{3}\right)+p_{0}+g z \frac{a^{3}}{\nu^{2}} . \tag{4.9}
\end{align*}
$$

The stresses (4.4) and (4.9) are now substituted in (4.1). After performing the integration in (4.1), it is found that, although all the coefficients $E_{2 s}$ and $F_{2 s}$ enter due to $\tau_{r \theta}$ and $\tau_{r r}$ considered separately, the end result depends only on $F_{2}$. The result is

$$
\begin{equation*}
F_{e}+W_{s}-W_{b}-4 \pi F_{2}=0, \tag{4.10}
\end{equation*}
$$

where $W_{b}=W_{b}^{\prime} / \rho \nu^{2}$ and $W_{b}^{\prime}$ is the weight of the displaced fluid, i.e. the buoyant force. The difference ( $W_{s}-W_{b}$ ) is the submerged weight of the particle, $W_{f}$. The term $4 \pi F_{2}$ is the viscous drag, $D$. Hence (4.10) may be written:

$$
\begin{equation*}
F_{e}+W_{f}=D=4 \pi F_{2} . \tag{4.11}
\end{equation*}
$$

At this point, using (4.8) for the pressure, it is a simple matter to compute the pressure drop per sphere, $\Delta p$, defined by

$$
\begin{equation*}
\Delta p=[p]_{\substack{R=R_{0} \\ z=\frac{1}{2} \beta}}-[p]_{z=-\frac{1}{2} \beta}^{\substack{R \\ \hline}} \tag{4.12}
\end{equation*}
$$

where $R_{0}$ is any radius. From the cylindrical co-ordinate form of $\psi$, (3.3), (3.4) and (2.17), it may be shown that most of the terms in (4.8) are periodic in $z$ and do not contribute to $\Delta p$. The end result is

$$
\begin{equation*}
\Delta p=-8 F_{2}-16 E_{2}-4 V \beta \tag{4.13}
\end{equation*}
$$

A mean pressure gradient, $p_{z}$, is defined by

$$
\begin{equation*}
p_{z}=\frac{\Delta p}{\beta}=-\frac{8}{\beta}\left(F_{2}+2 E_{2}\right)-4 V . \tag{4.14}
\end{equation*}
$$

For Poiseuille flow having the same discharge, $Q$, as the present case, but without any spheres present, the pressure gradient will be denoted $p_{z 0}$. In the present notation,

$$
\begin{equation*}
p_{z 0}=-4 V \tag{4.15}
\end{equation*}
$$

which is the last term of (4.14).

## 5. Numerical results

Computations were performed on the IBM 7094 computer system, first, to evaluate the coefficients $E_{2 s}$ and $F_{2 s}$. Successively larger truncations of (3.27) and (3.28) were solved, until the values of $E_{2 s}$ and $F_{2 s}$ did not change by more than one place in the fifth significant figure of $E_{2}$ and $F_{2}$.

Values of the stream function, $\psi$, were computed next, using a grid of at least $25 \times 25$ points, from which streamlines were drawn by interpolation. Velocity distributions on two cross-sections were computed using (2.9) and (3.24). The stress distribution on the spheres, drag and pressure drop were computed using (4.4), (4.9), (4.11) and (4.13). Computations were carried out for arbitrary sphere velocity, $U$, and arbitrary discharge, $Q=\frac{1}{2} \pi V$. The results are reported for each of the three cases (i) $U \neq 0, V=0$, (ii) $U=0, V \neq 0$, (iii) zero drag.

Figures 2 and 3 show a typical set of streamlines for the case of $\beta=1 \cdot 5$, $\lambda=0.5$. Figure $2(a)$ shows streamlines for the case of spheres moving in a fixed tube with zero discharge. Streamlines for the same flow, but viewed from the relative co-ordinate system fixed to the spheres, are shown in figure 2(b). Figure $2(c)$ shows streamlines for the case of flow with discharge $Q=\frac{1}{2} \pi V$ past fixed spheres in a fixed tube.

The streamline patterns in figure 2 are similar to those computed by Haberman \& Sayre (1958) for a single sphere in an infinitely long tube. One new feature is the presence in figure $2(a)$ of a layer near the tube wall, which moves continuously in the negative $z$-direction. No such layer is present in the case of a single sphere.

Figure 3 shows streamlines for the case of zero drag on the spheres, which results when a mixture of neutrally bouyant spheres and fluid is pumped through the tube by a pressure gradient. Figure $3(a)$ shows the streamlines relative to the fixed tube wall. Figure $3(b)$ shows the same flow viewed in relative co-ordinates, which move with the spheres. The most striking feature of these curves is that, in figure $3(a)$, all streamlines are nearly straight lines. Since the spheres are rigid, this means that all the fluid within a cylinder having the same radius as that of the spheres moves almost as a rigid body. This pattern is closely approximated by the flow sometimes referred to as a 'stacked coins' flow (Whitmore 1967), in which a continuous, neutrally buoyant stack of disks moves along the axis of the tube surrounded by a concentric layer of fluid.

Figure 3(b) shows that the flow between spheres, when viewed in a co-ordinate system fixed to them, is a vortex motion similar to that computed by Brandt \& Bugliarello (1965) in the spaces between disks, which extend to the full diameter of the tube.
The drag on each sphere is given by (4.11), in which the coefficient $F_{2}$ is in general of the form $\quad F_{2}=F_{2 U} U+F_{2 V} V$,
where $F_{2 U}$ and $F_{2 V}$ are coefficients computed when $E_{2 s}$ and $F_{2 s}$ are found. In dimensional form, the drag $D^{\prime}$ is
where

$$
\begin{align*}
& D^{\prime}=6 \pi \mu \lambda a\left(-K_{U} U^{\prime}+K_{V} V^{\prime}\right)  \tag{5.2}\\
& K_{U}=\frac{-2}{3 \lambda} F_{2 U}, \quad K_{V}=\frac{2}{3 \lambda} F_{2 V} \tag{5.3}
\end{align*}
$$



Figure 2. Streamlines and velocity profiles for $\beta=1 \cdot 5, \lambda=0 \cdot 5$. (a) Spheres moving ( $U \neq 0, V=0$ ). Streamlines shown in co-ordinates fixed to the cylinder. (b) Spheres moving ( $U \neq 0, V=0$ ). Streamlines shown in co-ordinates moving with the spheres. (c) Spheres and cylinder fixed ( $U=0, V \neq 0$ ). Streamlines shown for discharge $Q=\frac{1}{2} \pi V$.

The drag on a single sphere moving with a velocity $U^{\prime}$ in an infinite fluid whose velocity is $V^{\prime}$ at infinity is Stokes's drag,

$$
\begin{equation*}
D^{\prime}=6 \pi \mu \lambda a\left(-U^{\prime}+V^{\prime}\right) \tag{5.4}
\end{equation*}
$$

$K_{U}$ and $K_{V}$ in (5.2) are normalized drag coefficients equal to unity for Stokes's drag.


Figure 3. Streamlines and velocity profiles for $\beta=\mathbf{1} \cdot 5, \lambda=0 \cdot 5$, zero drag case. (a) Streamlines shown in co-ordinates fixed to the cylinder. (b) Streamlines shown in co-ordinates moving with the spheres.

Table 1 gives the coefficients $K_{U}$ computed for a line of spheres, except the last column, which contains values for a single sphere given by Haberman \& Sayre (1958). The results of the present computations for $\beta=40$ agree with those of Haberman \& Sayre (which correspond to $\beta=\infty$ ) to three or four significant figures for most cases.

|  |  |  |  |  |  | Haberman |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\lambda$ | $\beta=2 \lambda$ | $\beta=1.0$ | $\beta=1.4$ | $\beta=2.0$ | $\beta=4.0$ | $\beta=40$ | $\&$ Sayre |
| 0.1 | 0.473 | 1.225 | 1.258 | 1.263 | 1.263 | 1.263 | 1.263 |
| 0.2 | 0.8721 | 1.559 | 1.660 | 1.679 | 1.680 | 1.680 | 1.680 |
| 0.3 | 1.531 | 2.096 | 2.317 | 2.368 | 2.370 | 2.370 | 2.371 |
| 0.4 | 2.719 | 3.075 | 3.463 | 3.585 | 3.591 | 3.592 | 3.596 |
| 0.5 | 5.042 | 5.042 | 5.675 | 5.929 | 5.947 | 5.949 | 5.970 |
| 0.6 | 10.14 | - | 10.56 | 11.04 | 11.09 | 11.10 | 11.14 |
| 0.7 | 23.60 | - | 23.60 | 24.52 | 24.67 | 24.70 | 24.96 |
| 0.8 | 72.69 | - | - | 73.90 | 74.60 | 74.97 | 73.56 |
| 0.9 | 430.0 | - | - | 440.0 | 460.0 | - | - |

Table 1. Drag coefficient, $K_{U}$

|  |  |  |  |  |  | Haberman |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\lambda$ | $\beta=2 \lambda$ | $\beta=1.0$ | $\beta=1.4$ | $\beta=2.0$ | $\beta=4.0$ | $\beta=40$ | $\&$ Sayre |
| 0.1 | 0.469 | 1.216 | 1.249 | 1.255 | 1.255 | 1.255 | 1.255 |
| 0.2 | 0.8433 | 1.517 | 1.616 | 1.634 | .1635 | 1.635 | 1.635 |
| 0.3 | 1.422 | 1.964 | 2.177 | 2.227 | 2.229 | 2.229 | 2.231 |
| 0.4 | 2.393 | 2.723 | 3.092 | 3.210 | 3.216 | 3.216 | 3.218 |
| 0.5 | 4.158 | 4.158 | 4.735 | 4.977 | 4.995 | 4.996 | 5.004 |
| 0.6 | 7.750 | - | 8.121 | 8.562 | 8.613 | 8.617 | 8.651 |
| 0.7 | 16.54 | - | 16.54 | 17.34 | 17.47 | 17.49 | 17.67 |
| 0.8 | 46.12 | - | - | 47.06 | 47.58 | 47.81 | 47.30 |
| 0.9 | 250.0 | - | - | 250.0 | 260.0 | - | - |

Table 2. Drag coefficient, $K_{V}$

| $\lambda$ | $\beta=2 \lambda$ | $\beta=1.0$ | $\beta=1.4$ | $\beta=2.0$ | $\beta=4.0$ | $\beta=40$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0.1 | 1.983 | 1.987 | 1.987 | 1.987 | 1.987 | 1.987 |
| 0.2 | 1.934 | 1.944 | 1.947 | 1.947 | 1.947 | 1.947 |
| 0.3 | 1.857 | 1.873 | 1.879 | 1.881 | 1.881 | 1.881 |
| 0.4 | 1.760 | 1.771 | 1.786 | 1.791 | 1.791 | 1.791 |
| 0.5 | 1.649 | 1.649 | 1.669 | 1.679 | 1.680 | 1.680 |
| 0.6 | 1.528 | - | 1.538 | 1.551 | 1.553 | 1.553 |
| 0.7 | 1.401 | - | 1.401 | 1.414 | 1.416 | 1.416 |
| 0.8 | 1.271 | - | - | 1.274 | 1.276 | 1.276 |
| 0.9 | 1.14 | - | - | 1.14 | 1.14 | - |

Table 3. Ratio of $2 U / V$ for zero drag

| $\lambda$ | $\beta=2 \lambda$ | $\beta=1.0$ | $\beta=1.4$ | $\beta=2.0$ | $\beta=4.0$ | $\beta=40 \cdot 0$ | Happel \& Byrne |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  |  | $\beta=1.4$ | $\beta=40 \cdot 0$ |
| $0 \cdot 1$ | 0.558 | 1.450 | 1.488 | $1 \cdot 496$ | 1.496 | 1.496 | 1.489 | $1 \cdot 496$ |
| $0 \cdot 2$ | 1.959 | $3 \cdot 543$ | 3.776 | $3 \cdot 821$ | $3 \cdot 822$ | $3 \cdot 822$ | 3•774 | $3 \cdot 819$ |
| $0 \cdot 3$ | $4 \cdot 769$ | $6 \cdot 641$ | $7 \cdot 384$ | $7 \cdot 560$ | $7 \cdot 566$ | $7 \cdot 566$ | $7 \cdot 367$ | 7.543 |
| $0 \cdot 4$ | $10 \cdot 18$ | 11.65 | $13 \cdot 33$ | $13 \cdot 88$ | $13 \cdot 90$ | 13.91 | 13.26 | 13.79 |
| 0.5 | $20 \cdot 77$ | $20 \cdot 77$ | 23.94 | $25 \cdot 33$ | $25 \cdot 43$ | 25.44 | $23 \cdot 68$ | $24 \cdot 98$ |
| 0.6 | 43.18 | - | 45.54 | $48 \cdot 49$ | $48 \cdot 84$ | 48.86 | $44 \cdot 44$ | $47 \cdot 15$ |
| 0.7 | $98 \cdot 62$ | - | 98.62 | 104.5 | $105 \cdot 6$ | $105 \cdot 7$ | 93.21 | 98.94 |
| 0.8 | 283.9 | - | - | $291 \cdot 3$ | $295 \cdot 2$ | $296 \cdot 6$ | - | 263.1 |
| 0.9 | 1500 | - | $\cdots$ | 1500 | 1600 | - | - | - |

Table 4. Pressure drop coofficient, $P_{V}$

The effect of neighbouring spheres on the drag exerted on each sphere is small, if the spacing is greater than about one tube diameter ( $\beta>2$ ). The data in table 1 are plotted in figure 4 . It is seen that, the closer the spheres, the less becomes the drag per sphere. For large spheres, the effect of spacing is less pronounced. Table 2 and figure 5 show similar results for $K_{V}$; the qualitative results are the same as for $K_{U}$.

| Happel \& Byrne |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | ---: |
| $\lambda$ | $\beta=2 \lambda$ | $\beta=1.0$ | $\beta=1.4$ | $\beta=2.0$ | $\beta=4.0$ | $\beta=40$ | $\beta=1.4$ | $\beta=40$ |
| 0.1 | 0.563 | 1.460 | 1.499 | 1.506 | 1.506 | 1.506 | 1.499 | 1.506 |
| 0.2 | 2.024 | 3.640 | 3.877 | 3.922 | 3.924 | 3.924 | 3.877 | 3.923 |
| 0.3 | 5.118 | 7.069 | 7.837 | 8.018 | 8.024 | 8.025 | 7.839 | 8.021 |
| 0.4 | 11.49 | 13.07 | 14.84 | 15.41 | 15.44 | 15.44 | 14.85 | $15 \cdot 40$ |
| 0.5 | 24.95 | 24.95 | 28.41 | 29.86 | 29.97 | 29.98 | 28.38 | 29.74 |
| 0.6 | 55.80 | - | 58.47 | 61.65 | 62.01 | 62.04 | 57.80 | 60.72 |
| 0.7 | 138.9 | - | 138.9 | $145 \cdot 6$ | 146.8 | 147.0 | 133.0 | 139.7 |
| 0.8 | 442.8 | - | - | 451.7 | 456.8 | 459.0 | - | 412.7 |
| 0.9 | 2700 | - | - | 2700 | 2800 | - | - | - |

Table 5. Pressure drop coefficient, $P_{U}$

| $\lambda$ | $\beta=2 \lambda$ | $\beta=1.0$ | $\beta=1 \cdot 4$ | $\beta=2 \cdot 0$ | $\beta=4.0$ | $\beta=40 \cdot 0$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $0 \cdot 1$ | $0.730 \times 10^{-4}$ | $0 \cdot 1995 \times 10^{-4}$ | $0 \cdot 1427 \times 10^{-4}$ | $0 \cdot 1016 \times 10^{-4}$ | $0.501 \times 10^{-5}$ | $0.501 \times 10^{-6}$ |
| $0 \cdot 2$ | $0.1199 \times 10^{-2}$ | $0.6347 \times 10^{-3}$ | $0.453 \times 10^{-3}$ | $0.320 \times 10^{-3}$ | $0.160 \times 10^{-3}$ | $0 \cdot 166 \times 10^{-4}$ |
| 0.3 | $0.6177 \times 10^{-2}$ | $0.4617 \times 10^{-2}$ | $0.3451 \times 10^{-2}$ | $0.2438 \times 10^{-2}$ | $0.1220 \times 10^{-2}$ | $0 \cdot 1220 \times 10^{-3}$ |
| 0.4 | 0.01953 | $0 \cdot 01775$ | 0.01430 | $0.999 \times 10^{-2}$ | $0.5199 \times 10^{-2}$ | $0.5199 \times 10^{-3}$ |
| 0.5 | 0.04933 | $0 \cdot 04933$ | 0.04245 | 0.03235 | 0.01631 | 0.001631 |
| 0.6 | 0.112 | - | $0 \cdot 1039$ | $0 \cdot 0840$ | 0.0428 | $0 \cdot 00429$ |
| 0.7 | 0.234 | - | $0 \cdot 2339$ | $0 \cdot 1961$ | 0.103 | $0 \cdot 0103$ |
| 0.8 | 0.500 | - | - | $0 \cdot 455$ | $0 \cdot 242$ | $0 \cdot 0243$ |
| 0.9 | $1 \cdot 3$ | - | - | 1.2 | $0 \cdot 64$ | - |

Table 6. Pressure gradient coefficient, $G_{V 0}$, for zero drag

The above results are utilized for the case of zero drag for neutrally buoyant spheres by setting $D^{\prime}$ (5.2) equal to zero, and solving for $U$ in terms of $V$. The ratio $2 U / V$ is the ratio of the velocity of the spheres to the mean flow velocity of the mixture. Numerical values are given in table 3. Small spheres travel at nearly twice the mean velocity, but as $\lambda$ increases the sphere velocity approaches the mean velocity from above.

The pressure drop per sphere and the pressure gradient are given, in general, by (4.13) and (4.14), in which $F_{2}$ and $E_{2}$ are each linear in $U$ and in $V$. Hence, in dimensional form, the total pressure drop per sphere may be written

$$
\begin{equation*}
\Delta p^{\prime}=\frac{\mu}{a}\left(-V^{\prime} P_{V}+U^{\prime} P_{U}\right)-\frac{4 \mu V^{\prime}}{a} \beta+\rho g \beta a ; \tag{5.5}
\end{equation*}
$$

and the total pressure gradient in dimensional form may be written:

$$
\begin{equation*}
p_{z}^{\prime}=\frac{4 \mu}{a^{2}}\left(-V^{\prime} G_{V}+U^{\prime} G_{U}\right)-\frac{4 \mu V^{\prime}}{a^{2}}+\rho g \tag{5.6}
\end{equation*}
$$

where $P_{V}, P_{U}, G_{V}$ and $G_{U}$ are coefficients derived from the $E_{2}$ and $F_{2}$ values computed previously. The term ( $4 \mu V^{\prime} / a^{2}$ ) in (5.6) is the Poiseuille flow pressure gradient; likewise, the term $\left(4 \mu V^{\prime} \beta / a\right)$ in (5.5) is the pressure drop in a length $\beta a$ in Poiseuille flow. The terms involving $g$ are the hydrostatic effects. Hence, in (5.6), the coefficients $G_{V}$ and $G_{U}$ are a measure of the additional pressure gradient due to the presence of the spheres. Similarly, $P_{V}$ and $P_{U}$ are a measure of the additional pressure drop, due to the spheres, in a length $\beta a$.


Figure 4. Drag coefficient, $K_{U}$.

Table 4 gives the pressure drop coefficient $P_{V}$. The last two columns are the results computed from the following approximate formula, which is given by Happel \& Byrne (1954):

$$
\begin{equation*}
\Delta p_{a}^{\prime}=\frac{2 D^{\prime}}{\pi a^{2}}\left[1-\frac{2}{3} \lambda^{2}\right] . \tag{5.7}
\end{equation*}
$$

The last two columns of table 4 are computed using (5.7), with the drag $D^{\prime}$ based on the coefficients in table 2 . The approximate results are good for $\lambda \leqslant 0.5$.

Table 5 gives the computed values of $P_{U}$. Figures 6 and 7 show ( $1+P_{V}$ ) and $\left(1+P_{U}\right)$ as functions of $\lambda$. It may be seen that the qualitative trends in the pressure drop coefficients are the same as for the drag coefficients discussed above.

In the case of zero drag, $U$ depends on $V$, as shown in table 3. Hence in dimensional form the pressure gradient may be written

$$
\begin{equation*}
p_{z}^{\prime}=\frac{4 \mu}{a^{2}}\left(-V^{\prime} G_{V 0}\right)-\frac{4 \mu V^{\prime}}{a^{2}}+\rho g \tag{5.8}
\end{equation*}
$$

where $G_{V 0}$ is a coefficient which is a measure of the additional pressure gradient due to the presence of the spheres compared to that for Poiseuille flow having the same discharge. Numerical values of $G_{V 0}$ are given in table 6 and plotted


Figure 5. Drag coefficient, $K_{\boldsymbol{V}}$.
in figure 8. These show that the additional pressure gradient is generally small, and is less than $50 \%$ over that for Poiseuille flow for $\lambda$ less than 0.8 . The upper curve in figure 8 shows the comparable coefficient, $G_{V 0}$, for the stacked coins model (Whitmore 1967), which in the present termnology is

$$
\begin{equation*}
G_{V 0}=\frac{\lambda^{4}}{1-\lambda^{4}} . \tag{5.9}
\end{equation*}
$$

The results show that the additional pressure drop for the case of spheres touching is about $75 \%$ of that for the stacked coins model, when the coins and spheres have the same diameter.

The additional pressure drop per sphere in the zero drag case is

$$
\begin{equation*}
\Delta p^{\prime}=-\frac{\mu}{a} V^{\prime} P_{V 0} \tag{5.10}
\end{equation*}
$$

where values of the coefficient $P_{V 0}$ are given in table 7. Again, there is little interaction between spheres for $\beta>2$.


Figure 6. Pressure drop coefficient, $1+P_{V}$.

|  | $\beta=2 \lambda$ | $\beta=\mathrm{I} \cdot 0$ | $\beta=\mathbf{1 . 4}$ | $\beta=2.0$ | $\beta=4.0$ | $\beta=40.0$ |
| :---: | :--- | :---: | :--- | :--- | :--- | :--- |
| $\lambda$ | $\beta=20^{-4}$ |  |  |  |  |  |
| 0.1 | $0.584 \times 10^{-4}$ | $0.789 \times 10^{-4}$ | $0.799 \times 10^{-4}$ | $0.800 \times 10^{-4}$ | $0.801 \times 10^{-4}$ | $0.801 \times 10^{-4}$ |
| 0.2 | $0.1919 \times 10^{-2}$ | $0.2359 \times 10^{-2}$ | $0.2559 \times 10^{-2}$ | $0.2562 \times 10^{-2}$ | $0.2562 \times 10^{-2}$ | $0.2562 \times 10^{-2}$ |
| 0.3 | 0.01482 | 0.01847 | 0.01932 | 0.01951 | 0.01951 | 0.01951 |
| 0.4 | 0.06249 | 0.07100 | 0.08010 | 0.08296 | 0.08318 | 0.08319 |
| 0.5 | 0.1973 | 0.1973 | 0.2377 | 0.2588 | 0.2609 | 0.2609 |
| 0.6 | 0.537 | - | 0.582 | 0.672 | 0.685 | 0.687 |
| 0.7 | 1.310 | - | 1.31 | 1.57 | 1.65 | 1.65 |
| 0.8 | 3.20 | - | - | 3.64 | 3.81 | 3.9 |
| 0.9 | 9.3 | - | - | 9.6 | 10.0 | - |

Tabla 7. Pressure drop coefficient, $P_{V 0}$, for zero drag

The results above have been developed and presented as if the particle velocity, $U$, and fluid discharge, $Q$ (or $V$ ), were given and the drag and pressure drop were to be found. The results above may also be used to find $U$ and $Q$ for particles of given size, spacing and density, assuming that the fluid properties, gravitational constant, pressure gradient and applied external force are also given. From the tabulated results and the particle size and spacing, the coefficients


Figure 7. Pressure drop coefficient, $1+\mathrm{P}_{U}$.
$K_{U}, K_{V}$ in (5.2), and $G_{U}, G_{V}$ in (5.6), may be determined. Then, equating the drag in (5.2) to the known applied external force and the weight of the particle in fluid, and substituting the known pressure gradient in (5.6), gives a pair of equations, which may be solved for $U$ and $V$.

## 6. Conclusion

With respect to the problem of capillary blood flow, the present study suggests two general conclusions. First, if the particles (red blood cells) are spaced more than one tube diameter apart, their interaction is negligible. Secondly, if the tube diameter is at least $20 \%$ greater than the particle diameter, the additional
pressure drop will not be greater than about $50 \%$ of the Poiseuille pressure drop of the suspending medium (blood plasma) alone. This is an interesting conclusion, because the viscosity of blood is generally at least twice that of the plasma alone in shear viscometer tests and in large tubes (Gregersen 1967). This corroborates experimental measurements in fine tubes (Haynes 1961), which show that the effective viscosity of blood is less in capillaries than in larger vessels.


Figure 8. Pressure gradient coefficient, $G_{V_{0}}$, for the case of zero drag. Solid curves are computed results for a line of spheres. Dotted curve is for the stacked coins model. Dashed curve ( $\beta=2 \lambda$ ) is for a line of spheres which are touching.

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